

Calibrating Stellar Models in the TESS Era

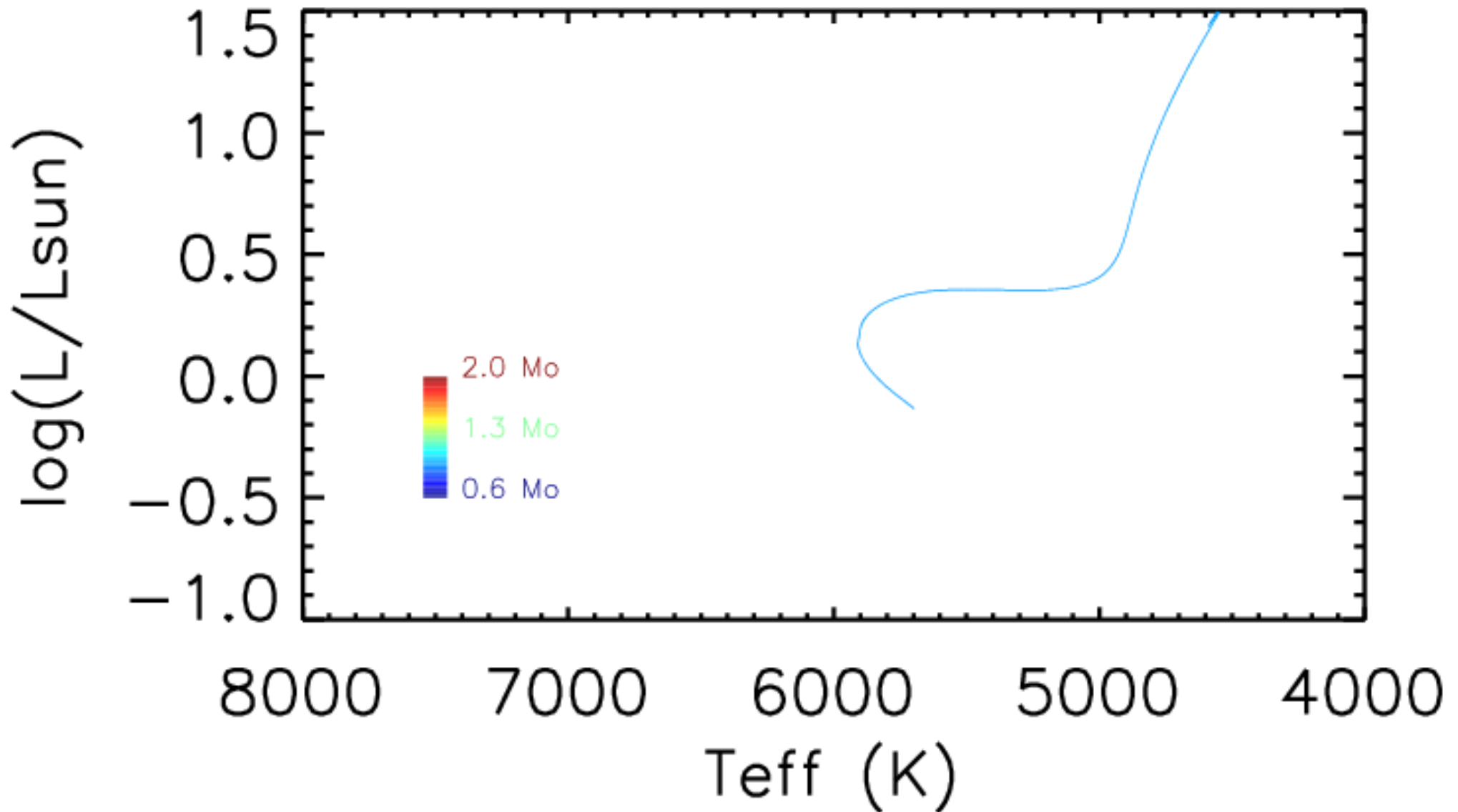
Jamie Tayar

With Zach Clayton and the
APOGEE-Kepler, APOGEE-TESS, & TESS Subgiant Collaborations

TESS Science Conference
July 30, 2019



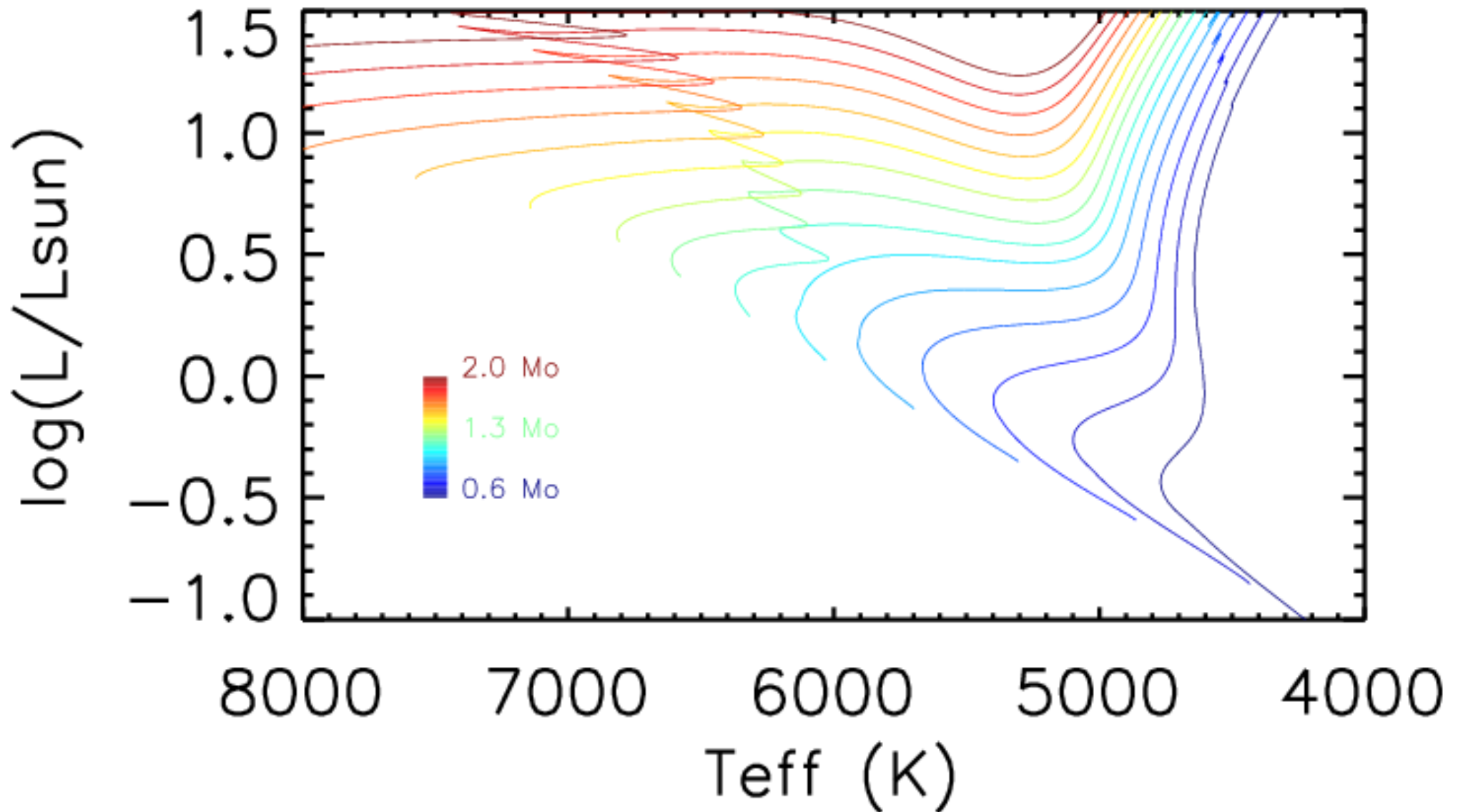
Stellar Model



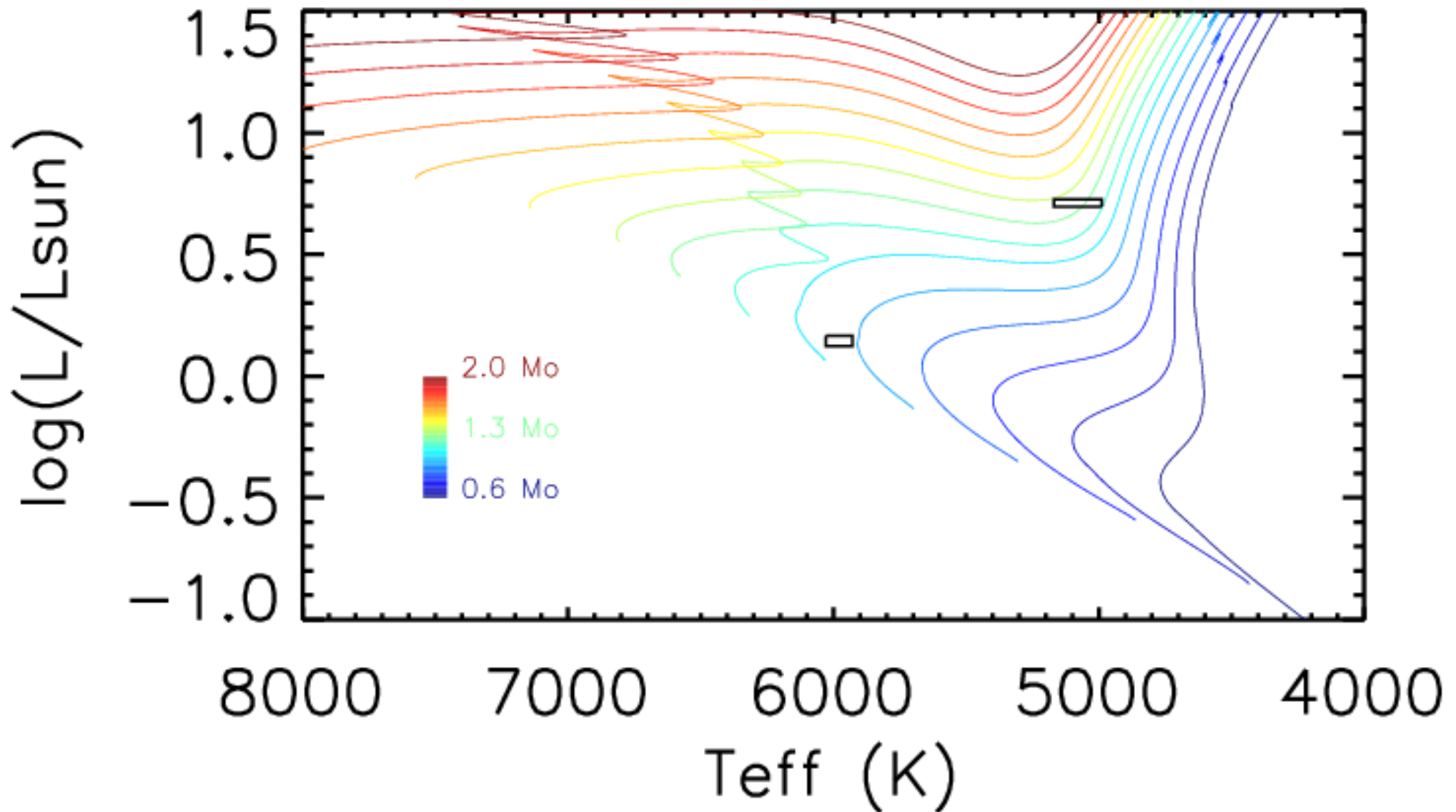
Model Physics Uncertain

- Convection
- Rotation
- Mixing
- Diffusion
- Atmosphere Boundary Condition
- Helium
- Solar composition
- ...

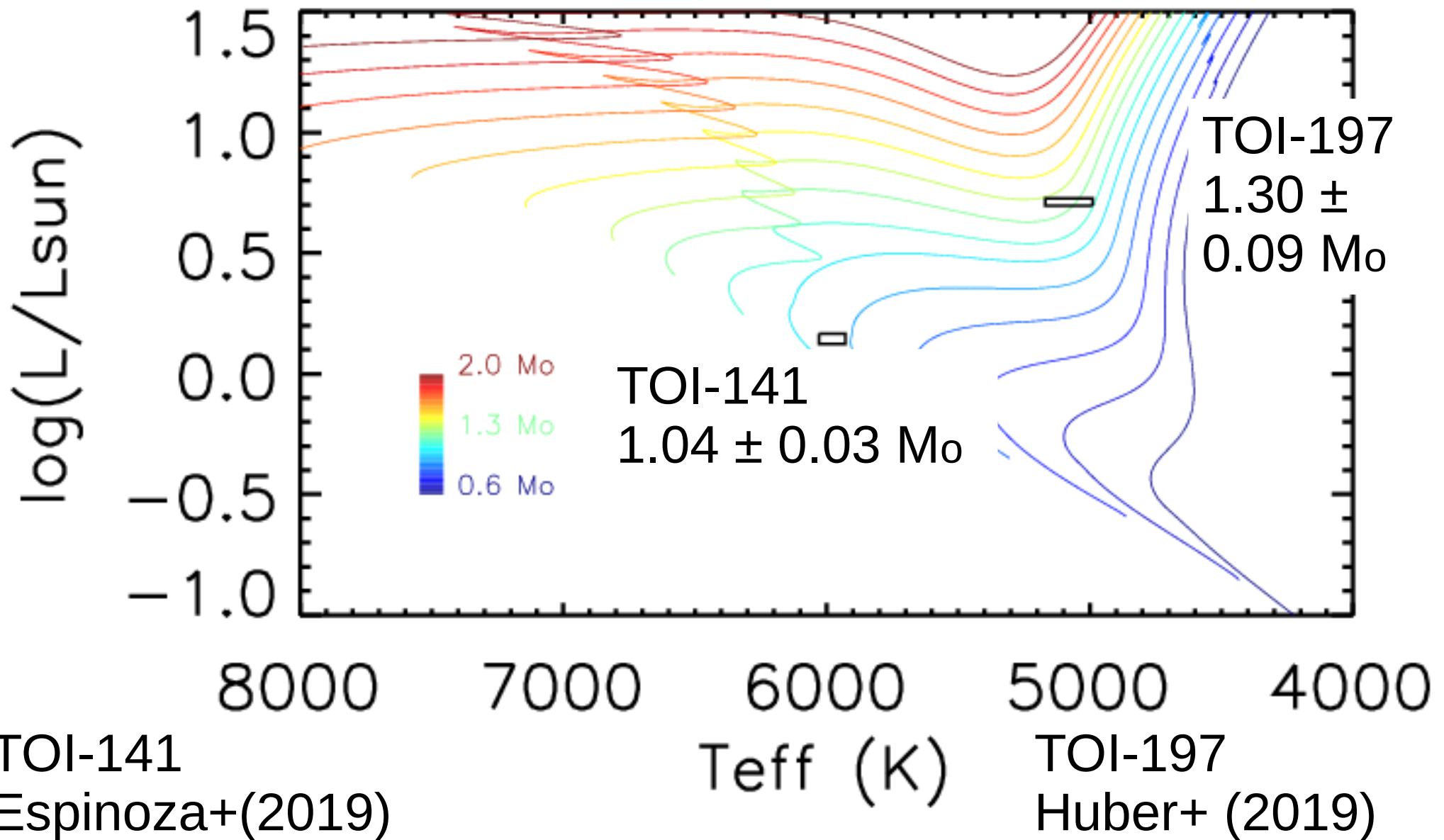
Stellar Model Grid



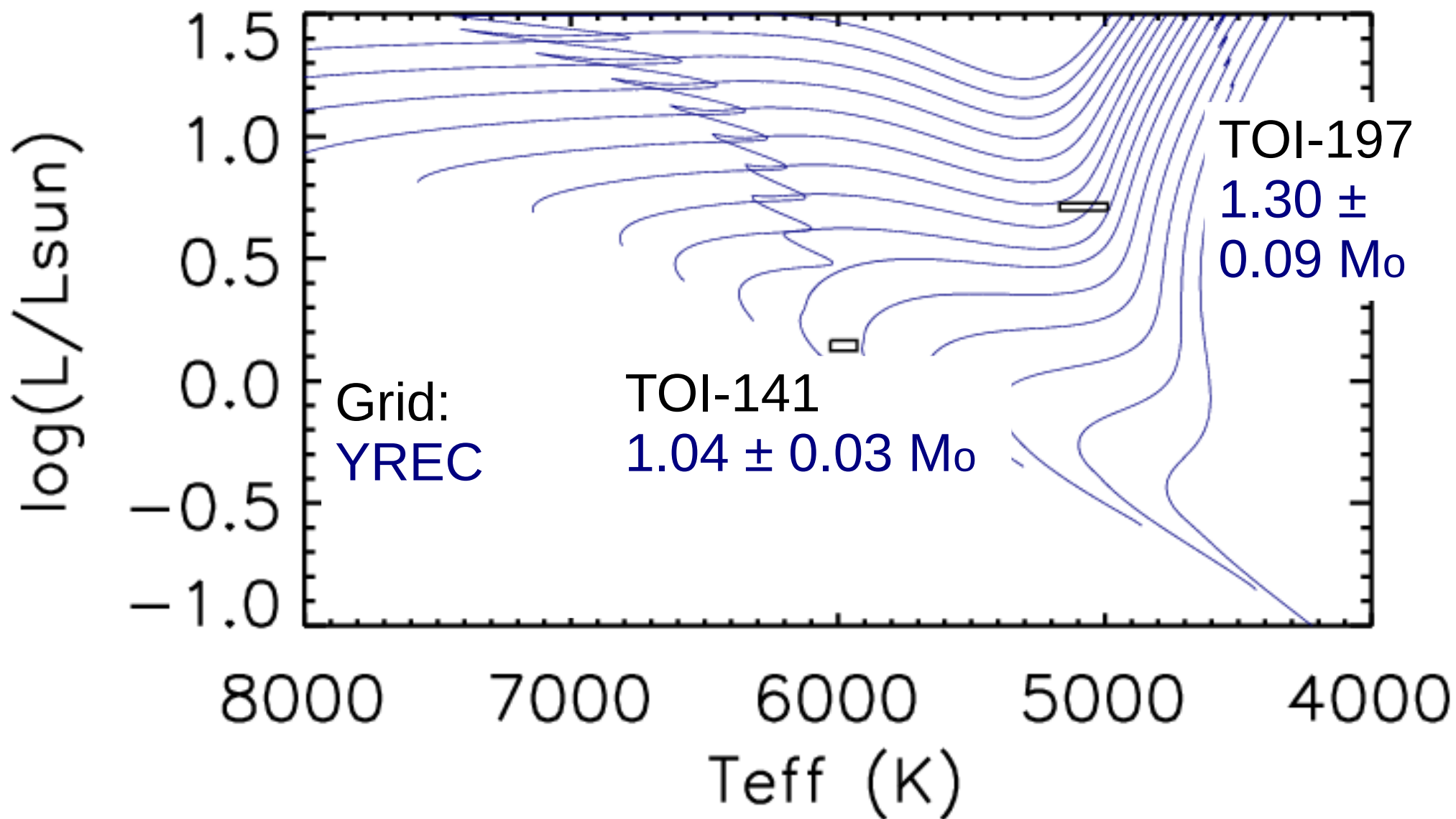
Stellar Model Grid



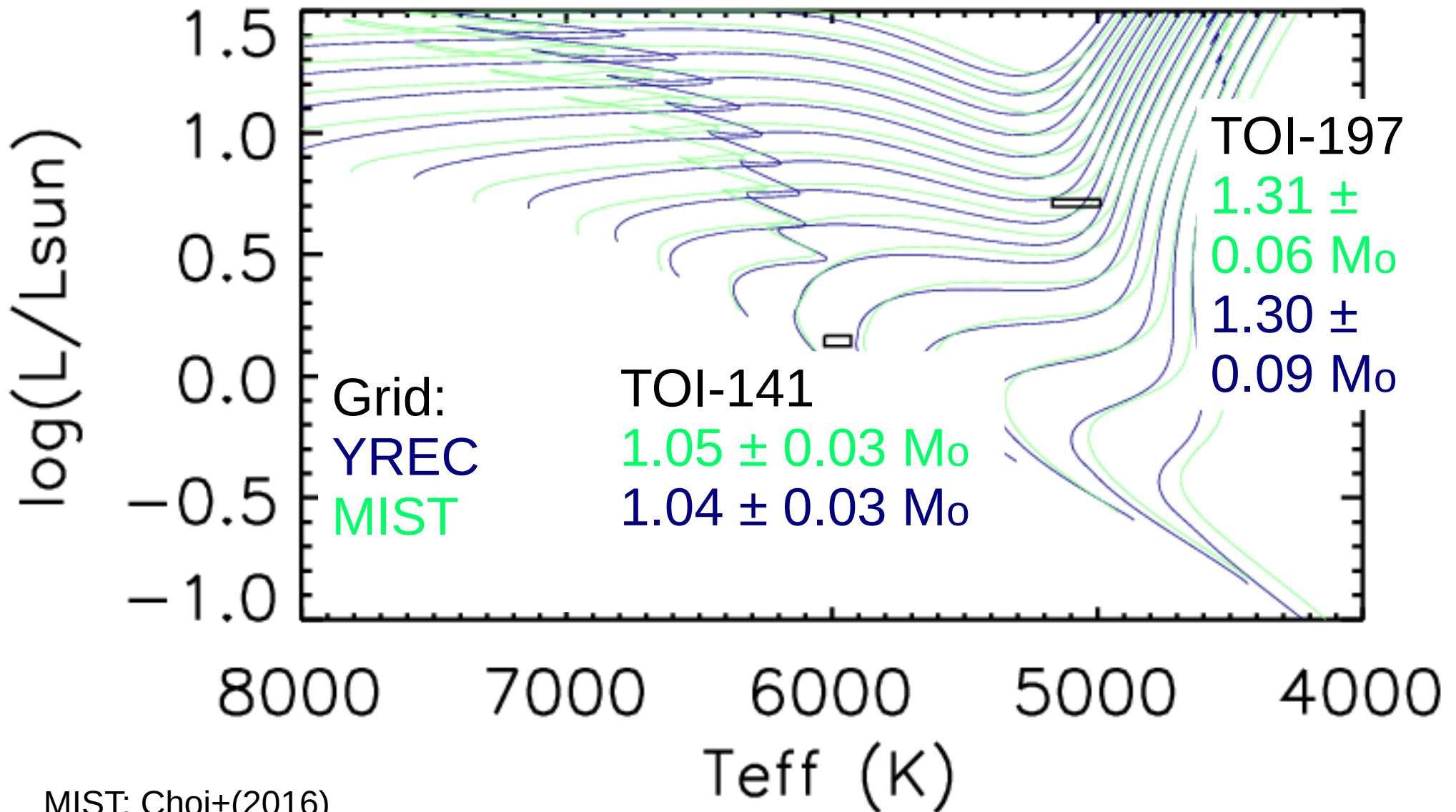
Estimate Stellar Properties



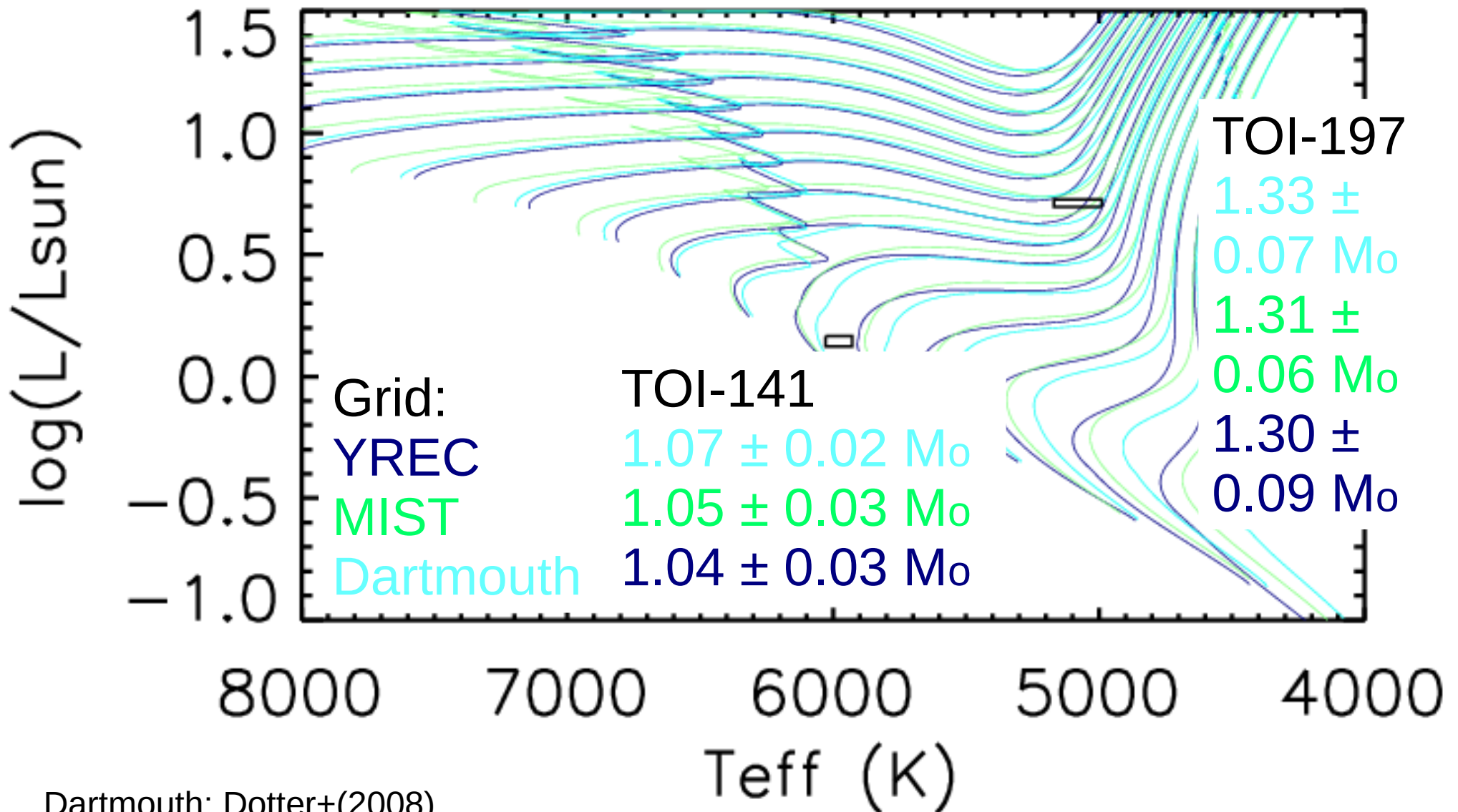
Impact of Model Choice: YREC



Impact of Model Choice: MIST

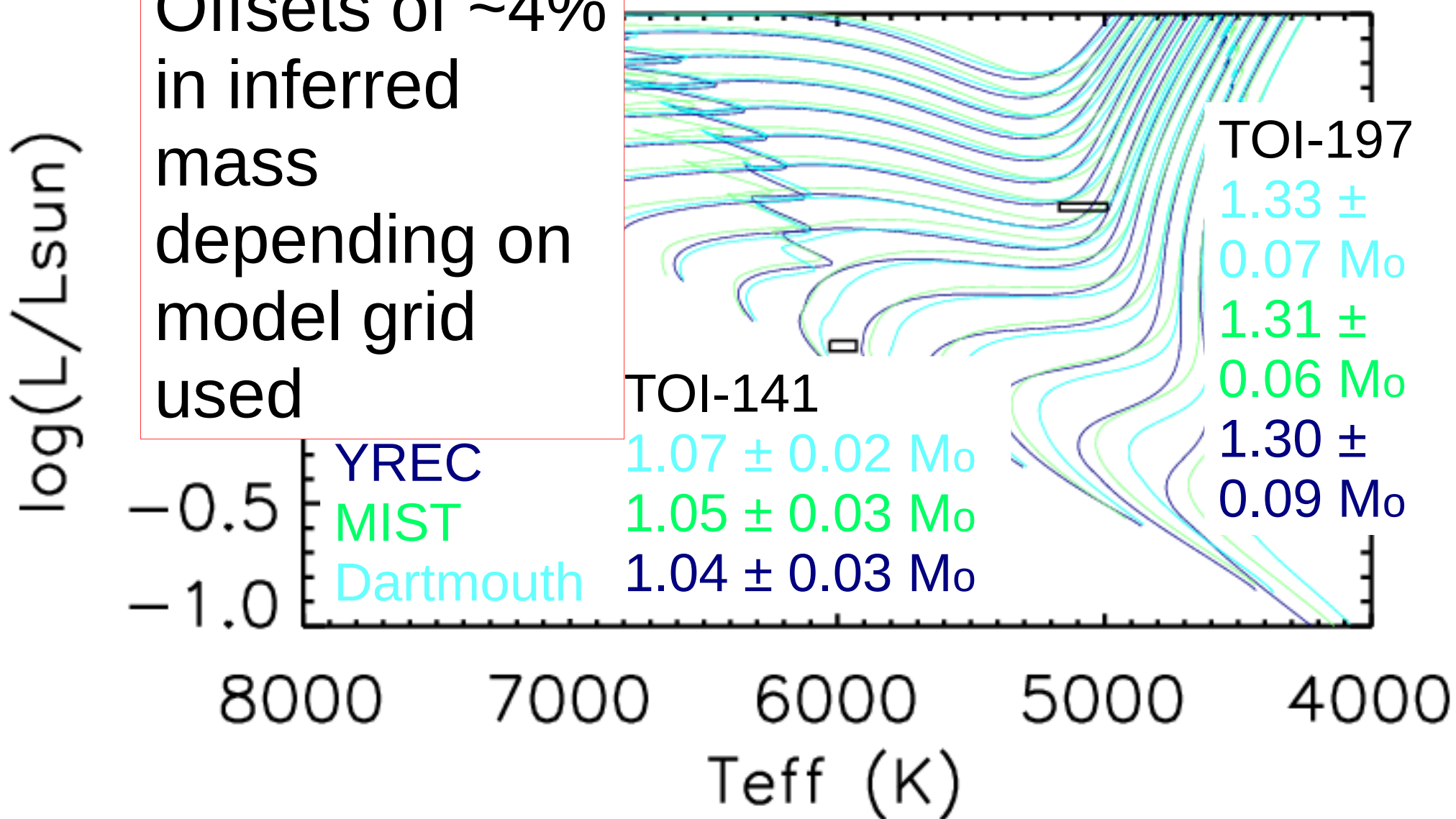


Impact of Model Choice: Dartmouth



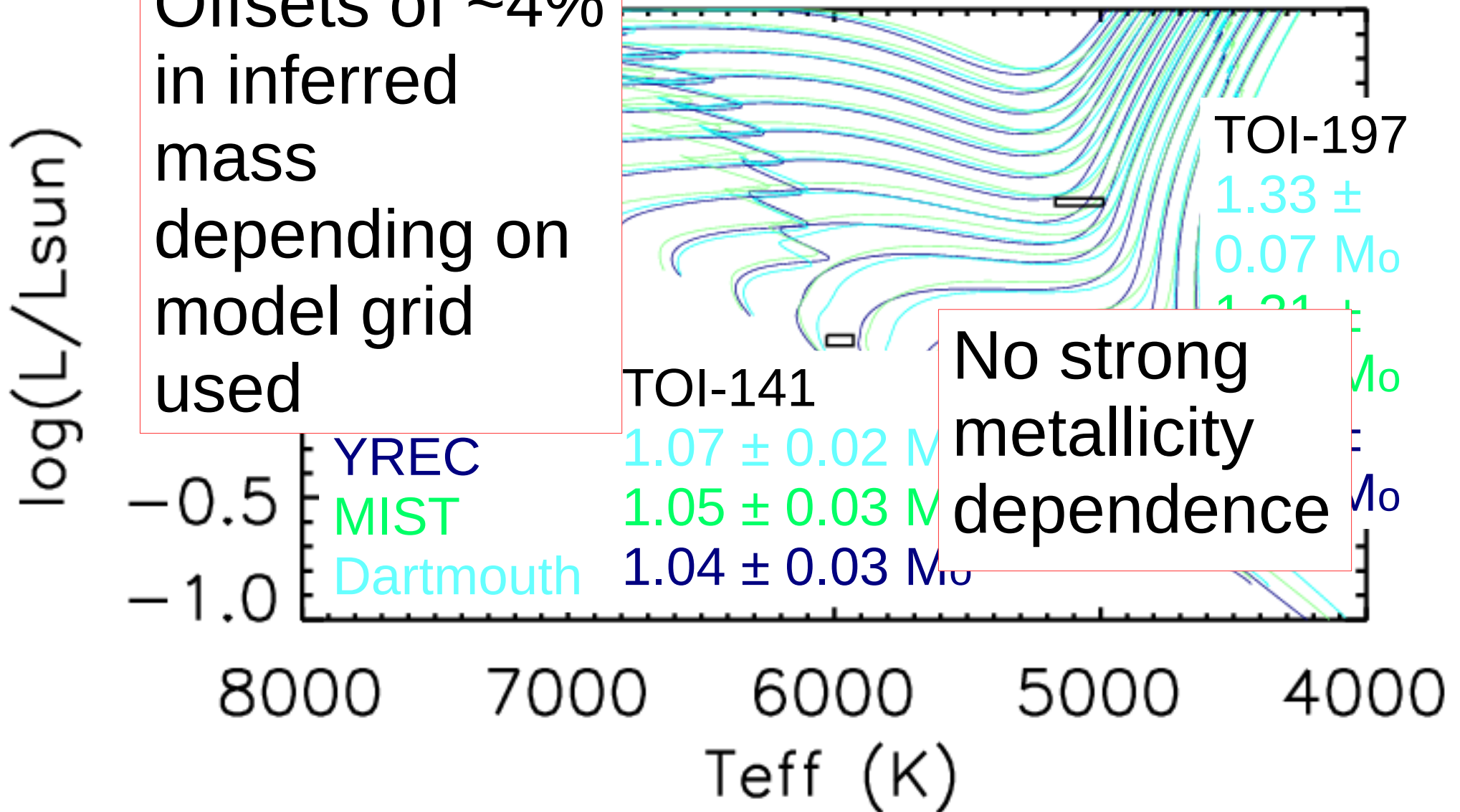
Model Choice Matters

Offsets of $\sim 4\%$ in inferred mass depending on model grid used



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Offsets of $\sim 4\%$ in inferred mass depending on model grid used

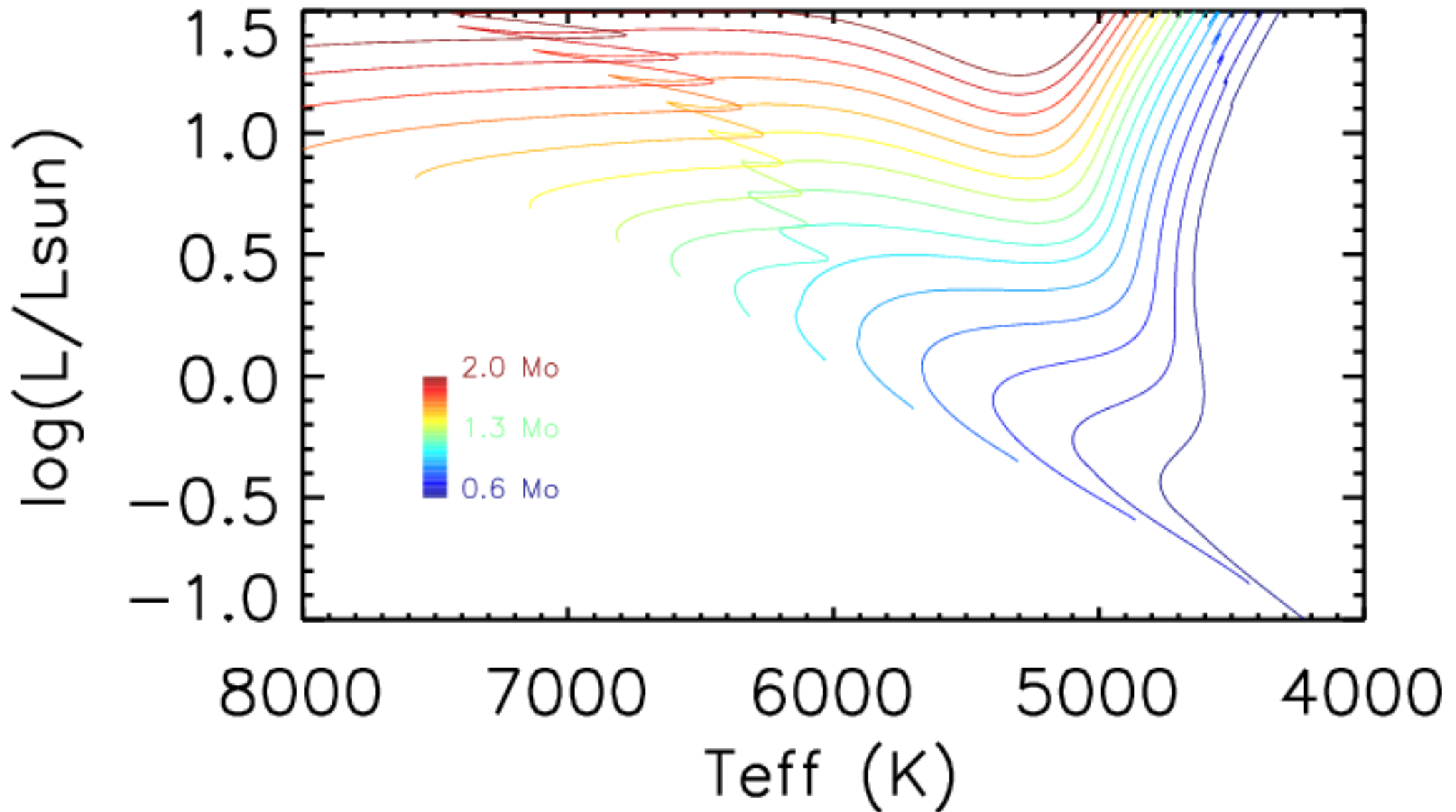


No strong metallicity dependence

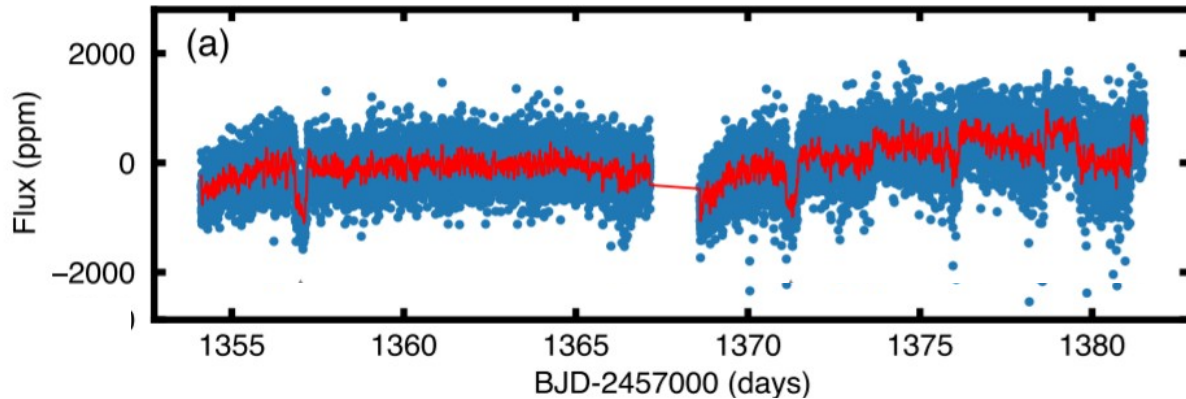
Take Away

When estimating stellar properties from models, either inflate the error bars to account for model systematics or compare to multiple models.

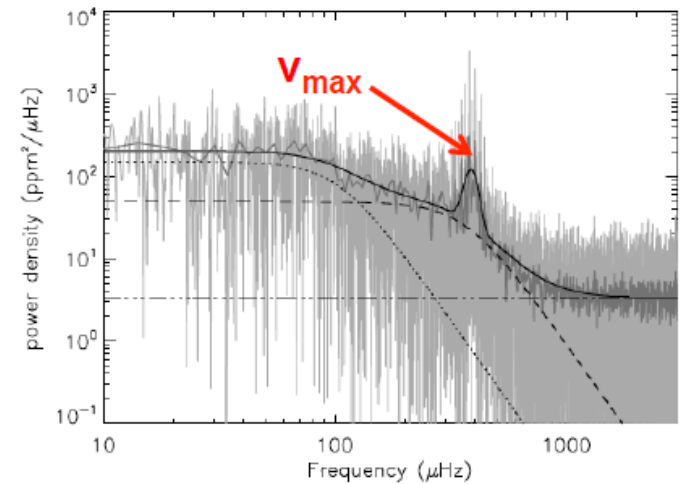
Stellar Model Grid



Asteroseismic Masses

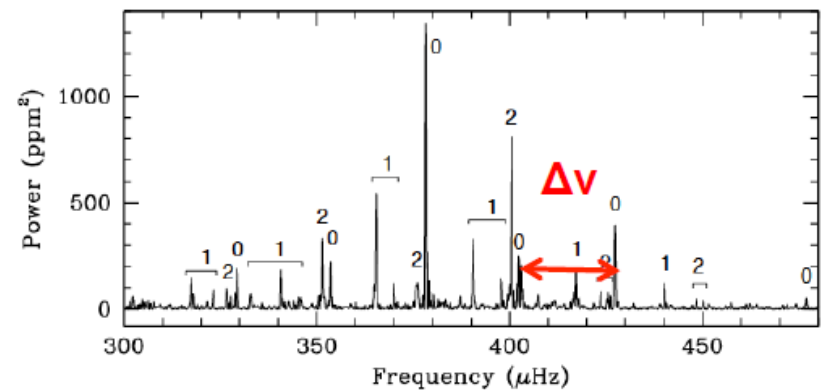


- Sample light curve from Huber+ (2019)



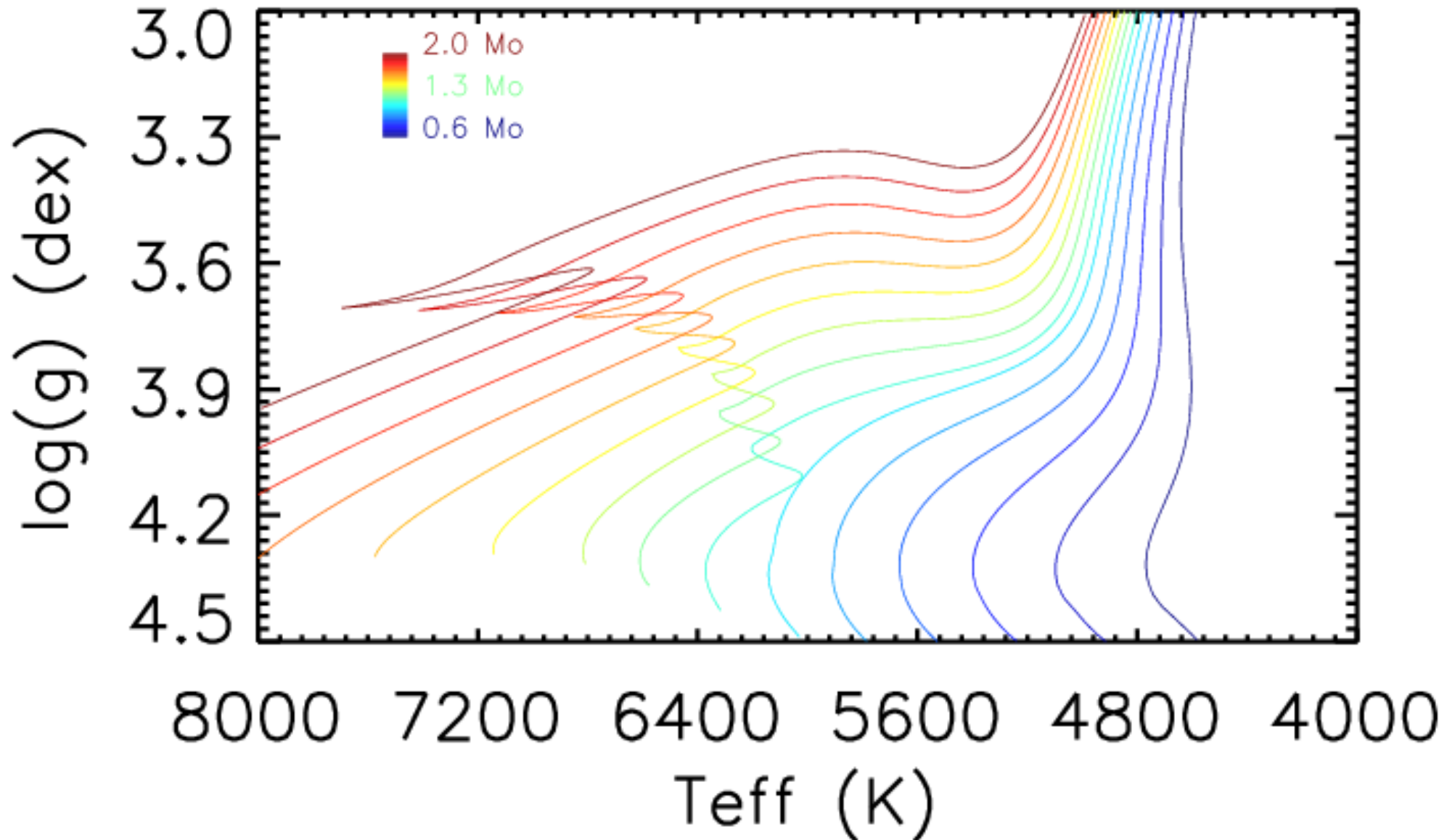
$$\frac{\Delta\nu}{\Delta\nu_{\odot}} = \sqrt{\frac{M/M_{\odot}}{(R/R_{\odot})^3}}$$

$$\frac{\nu_{\max}}{\nu_{\max,\odot}} = \frac{M/M_{\odot}}{(R/R_{\odot})^2 \sqrt{T_{\text{eff}}/T_{\text{eff},\odot}}}$$

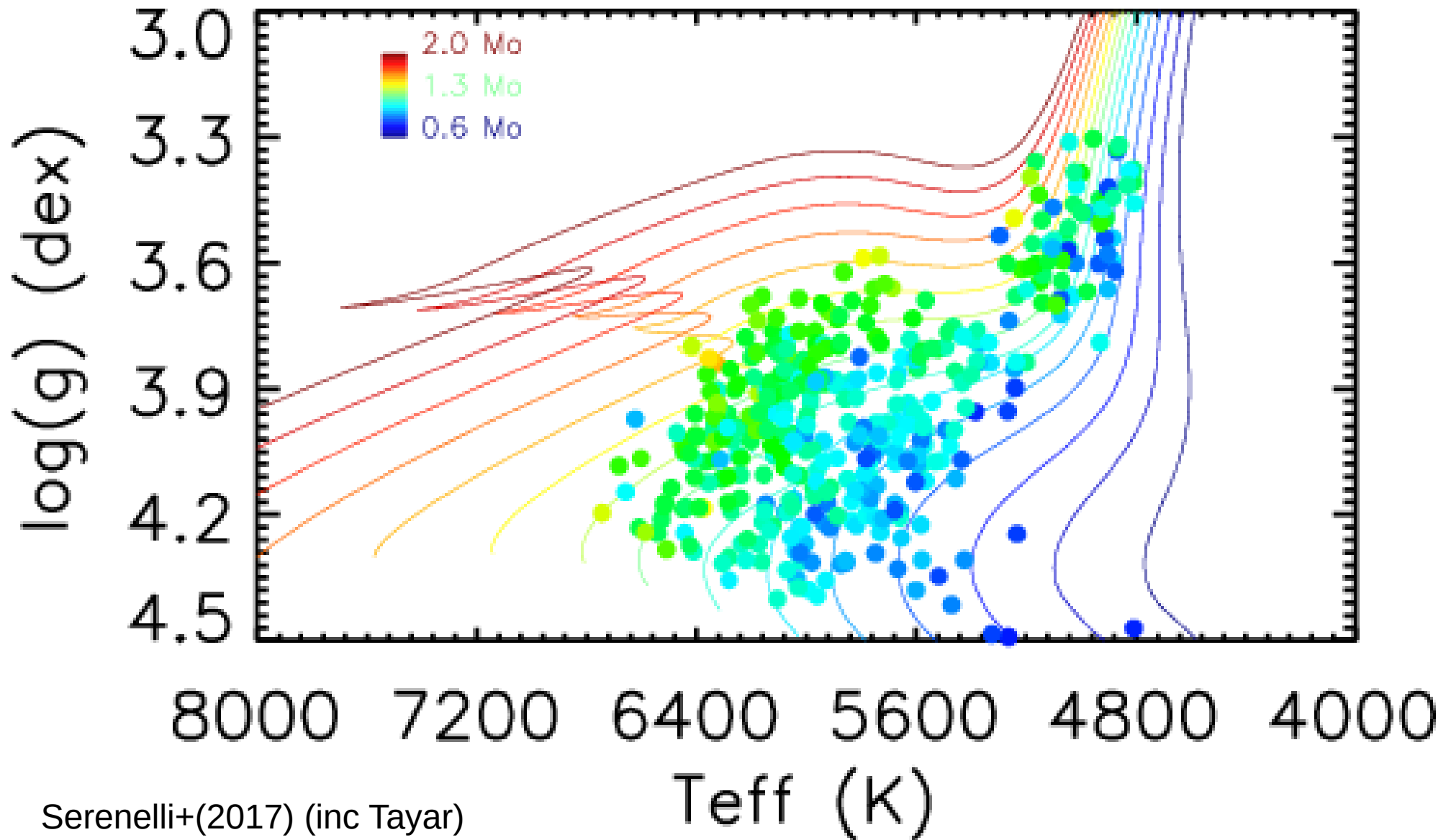


Di Mauro+ (2011)

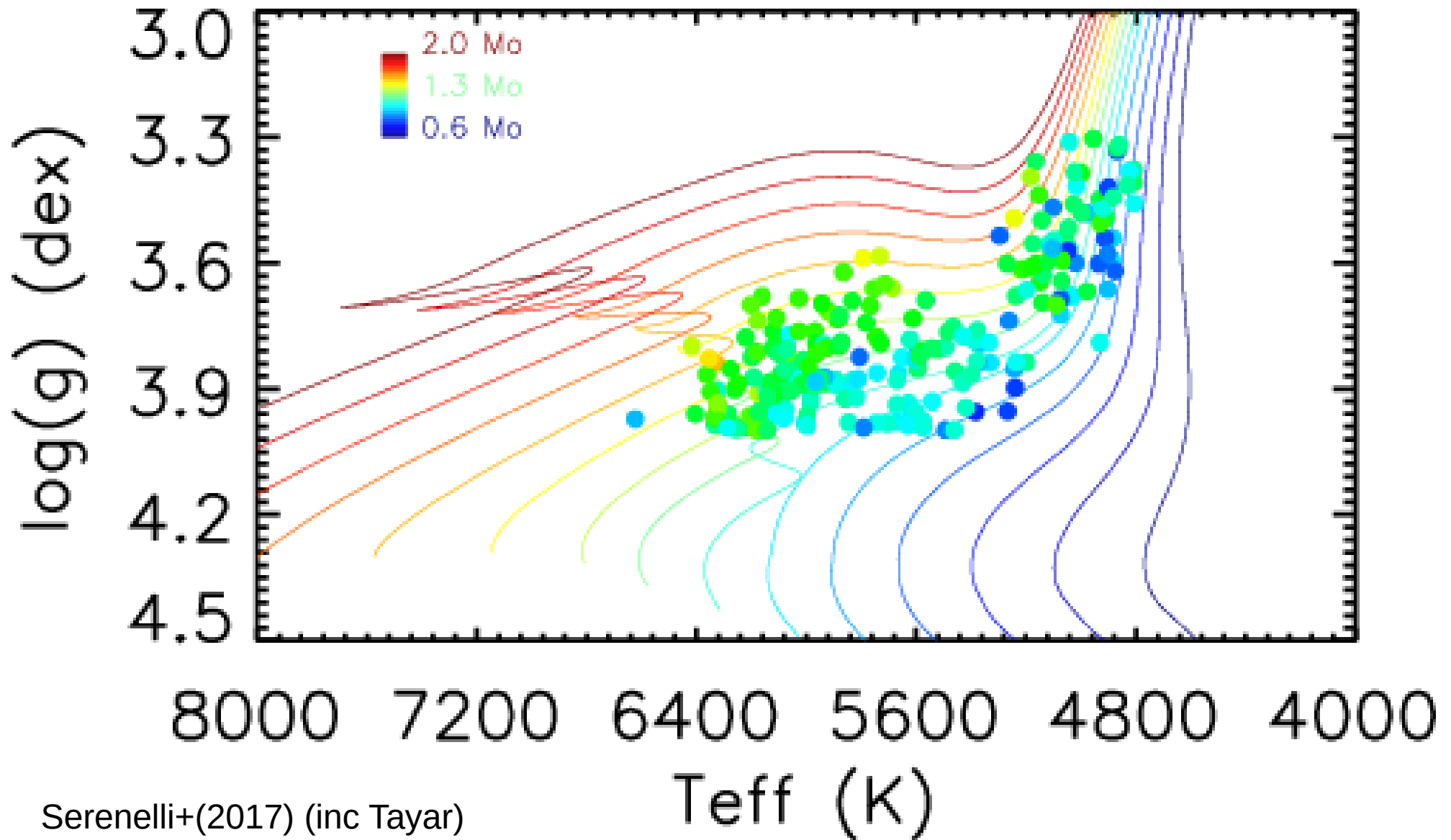
Stellar Models



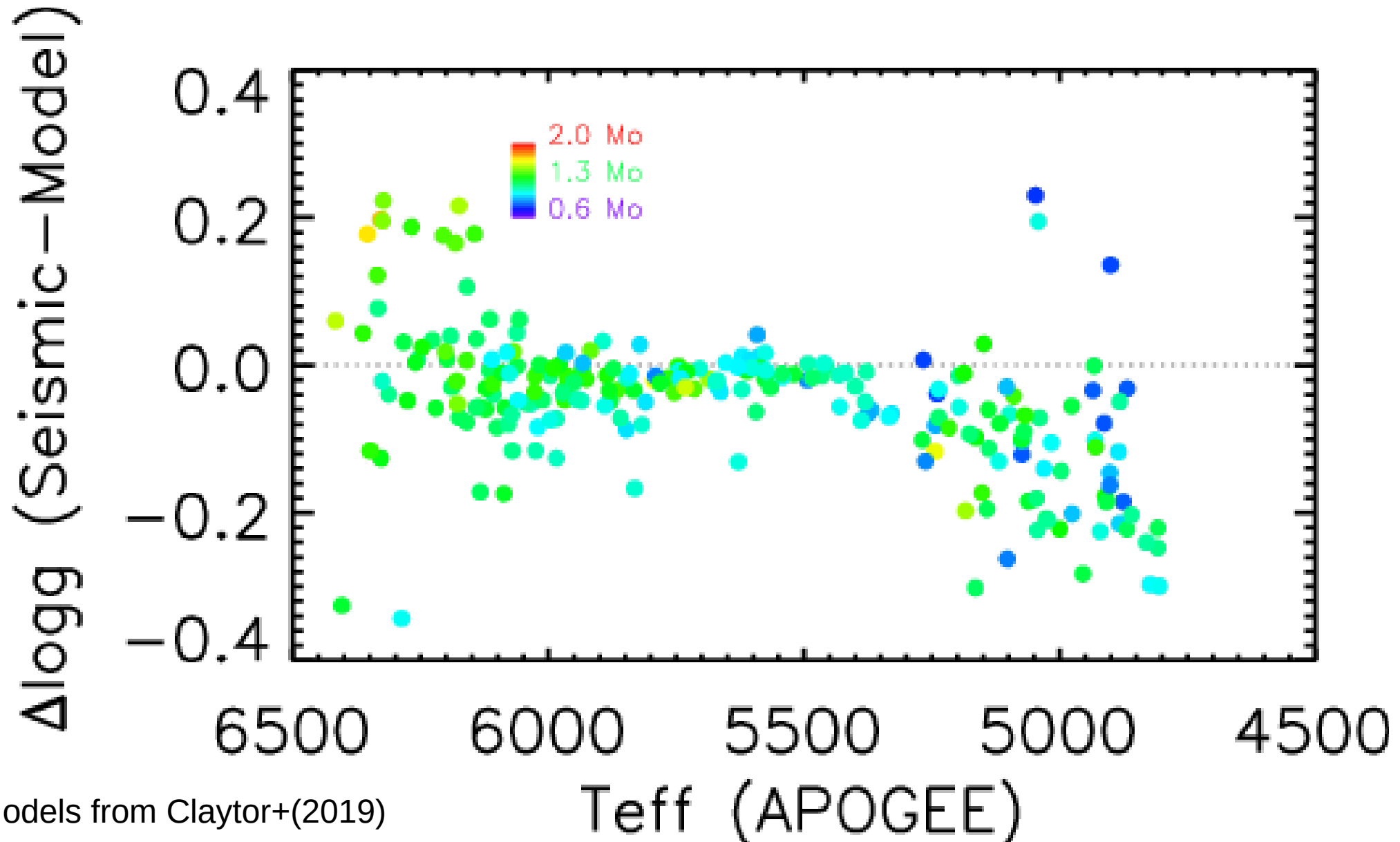
APOGEE-Kepler Subgiants



APOGEE-Kepler Subgiants

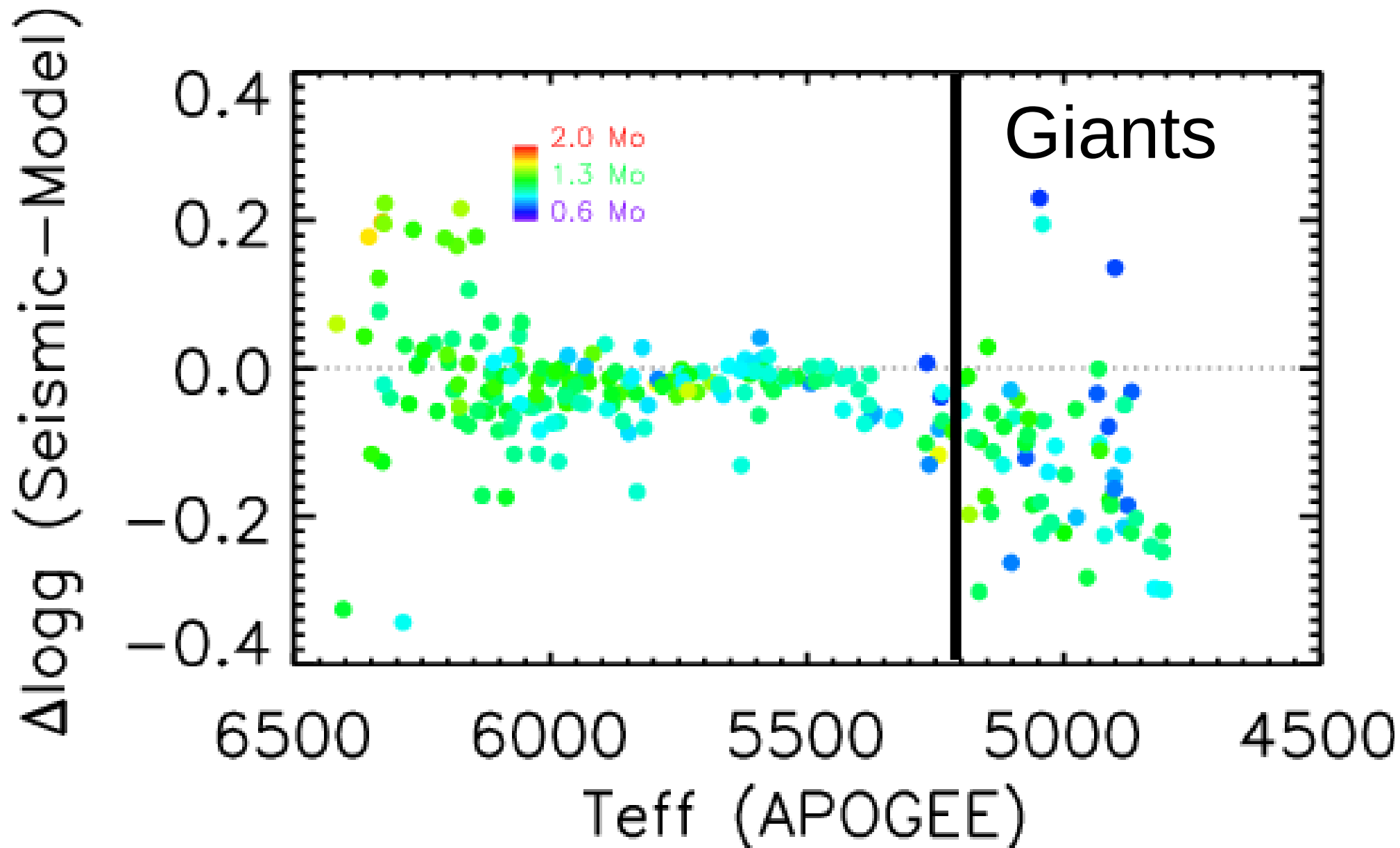


APOGEE-Kepler Subgiants

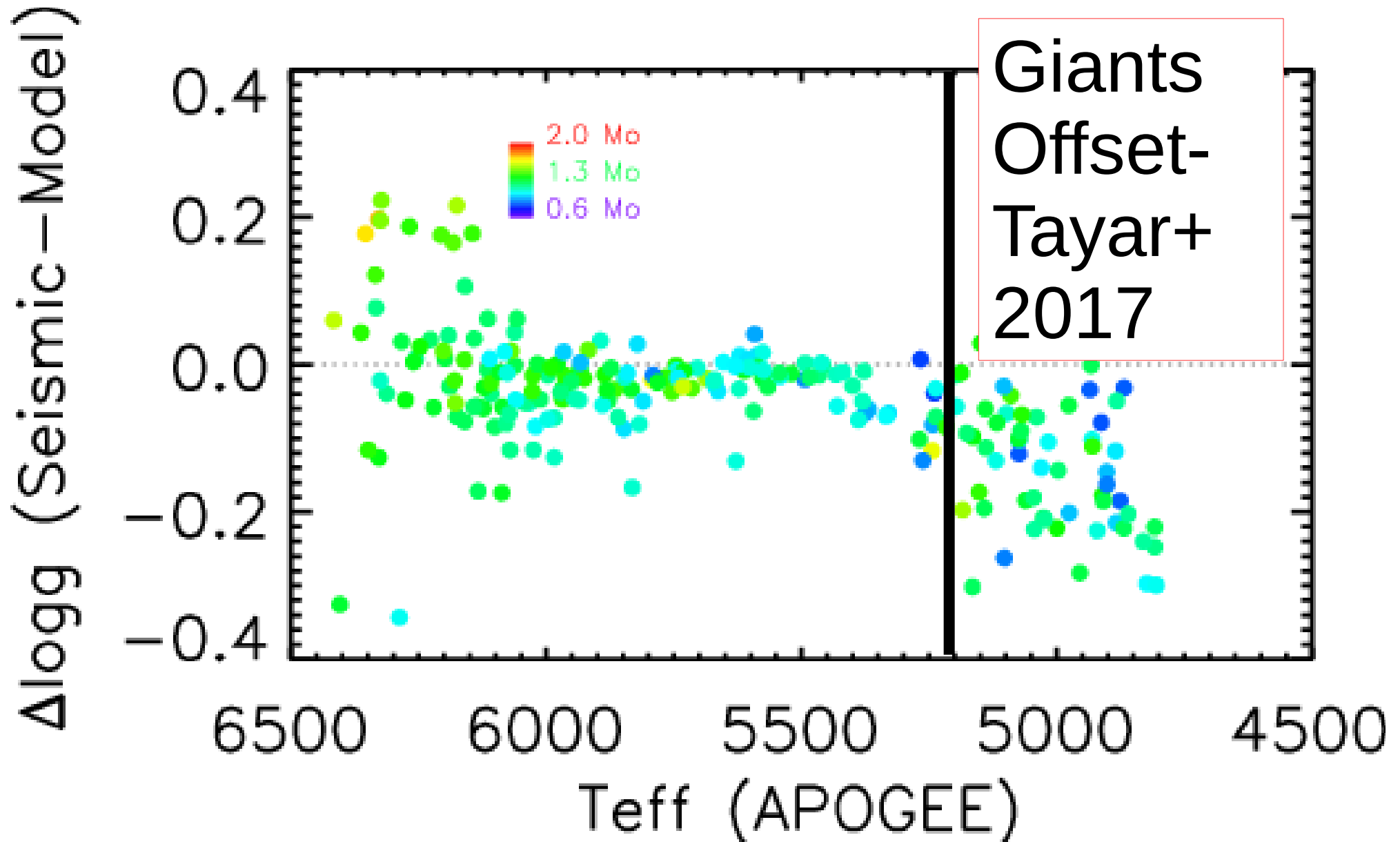


Models from Clayton+(2019)

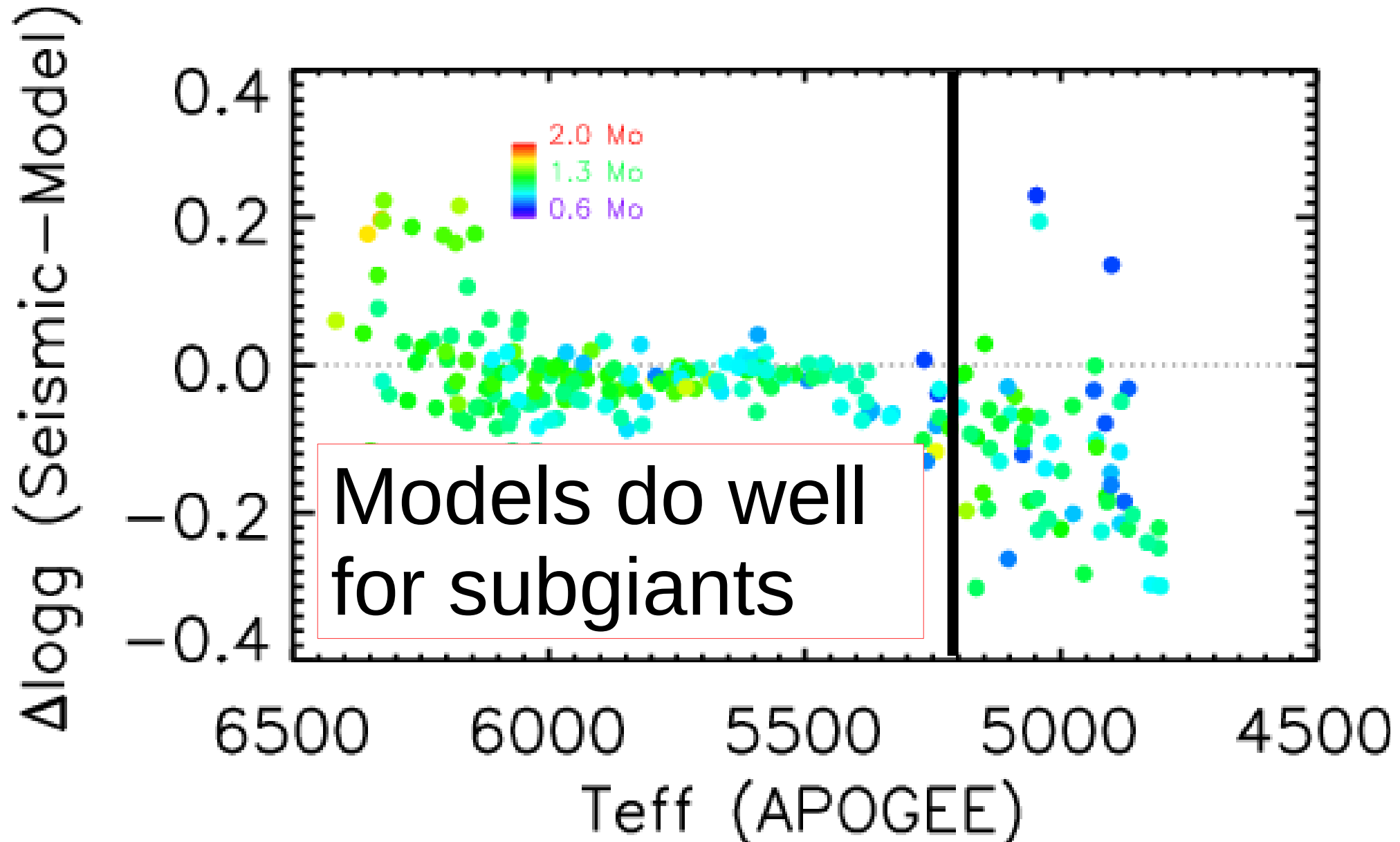
APOGEE-Kepler Subgiants



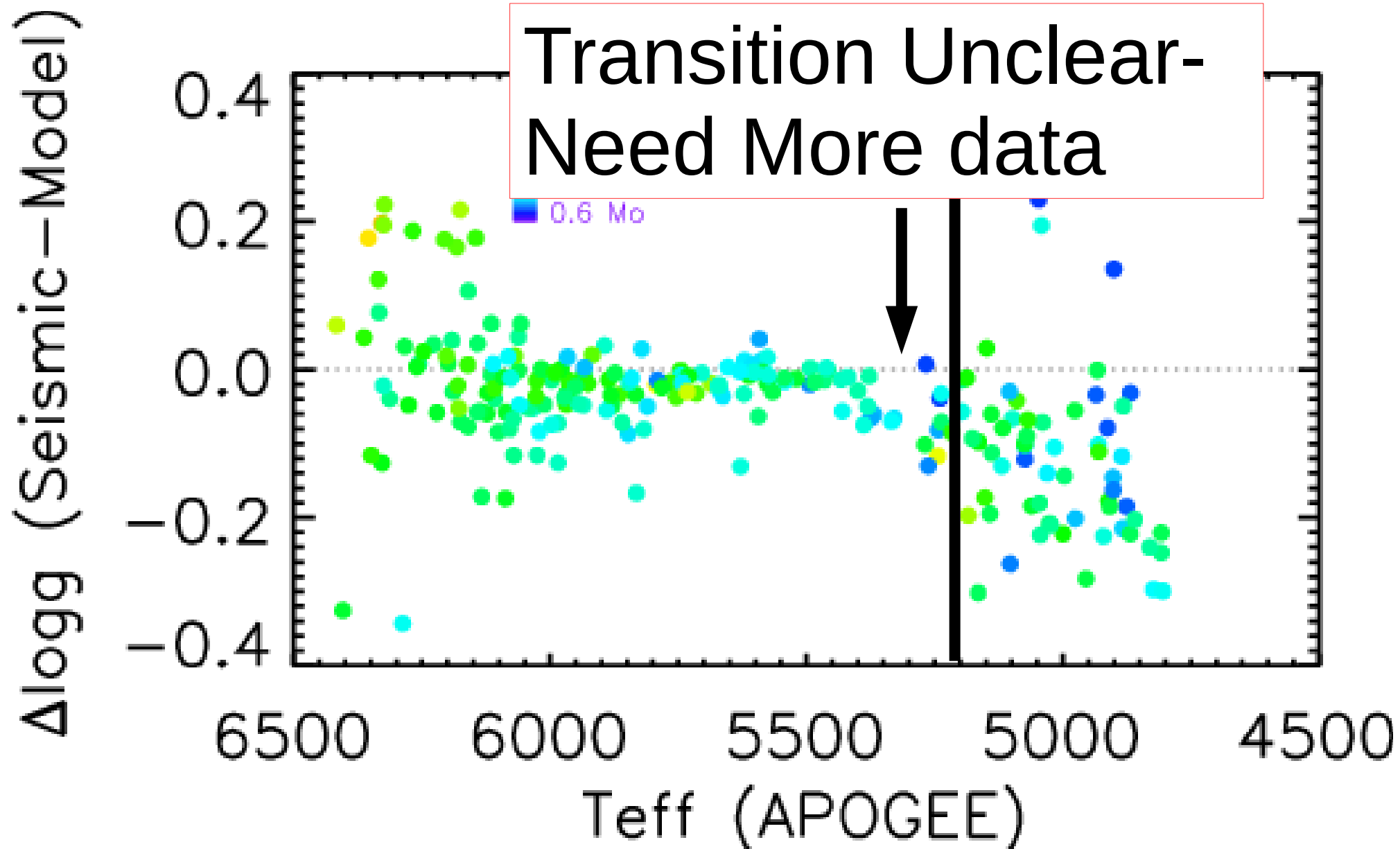
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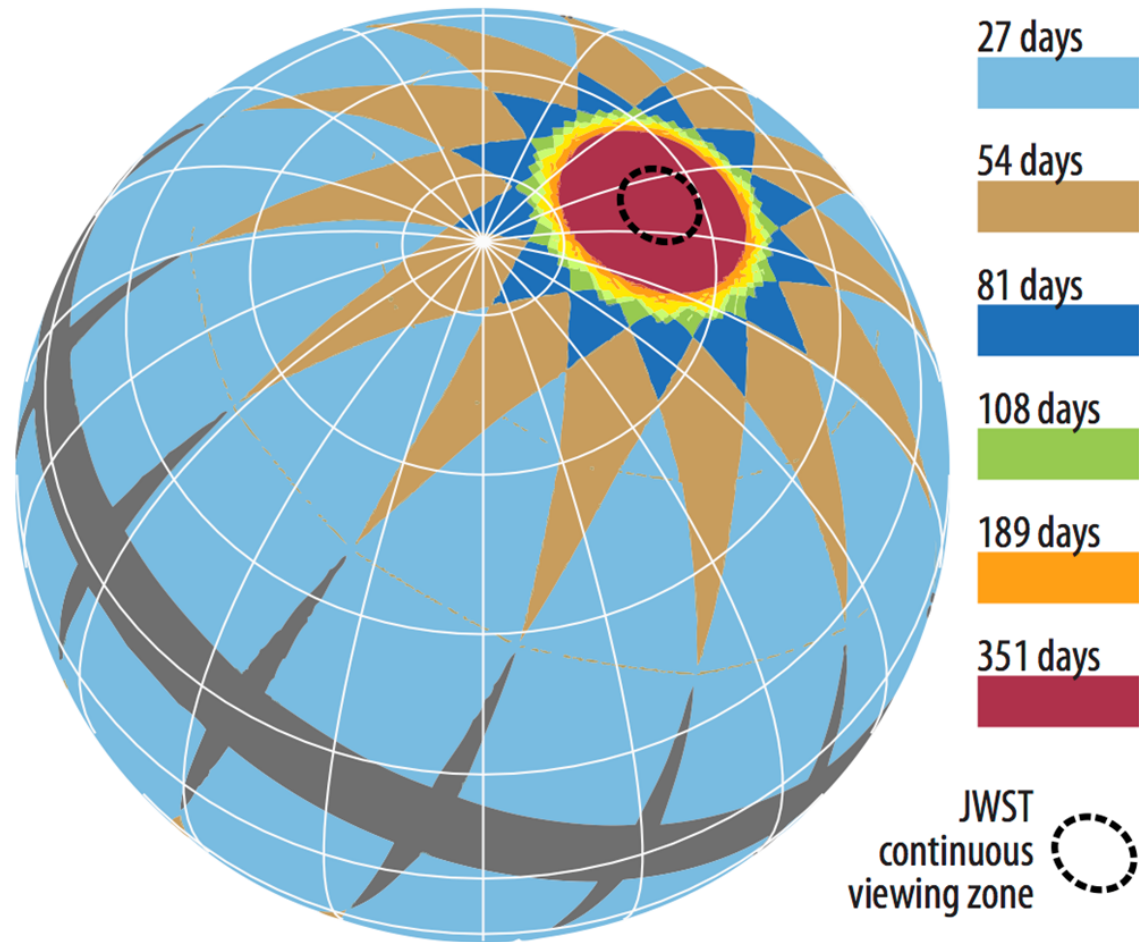


APOGEE-Kepler Subgiants

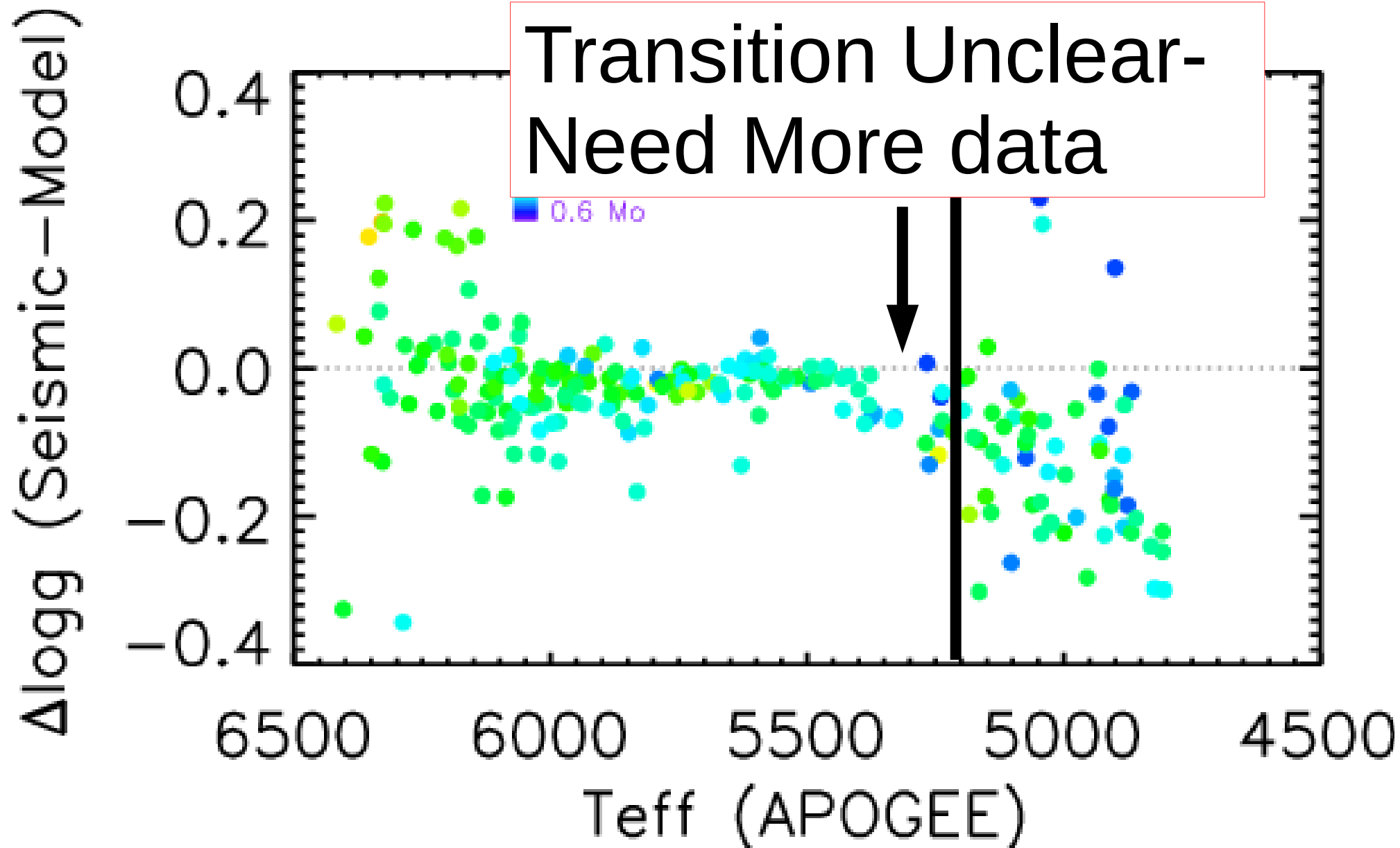


TESS

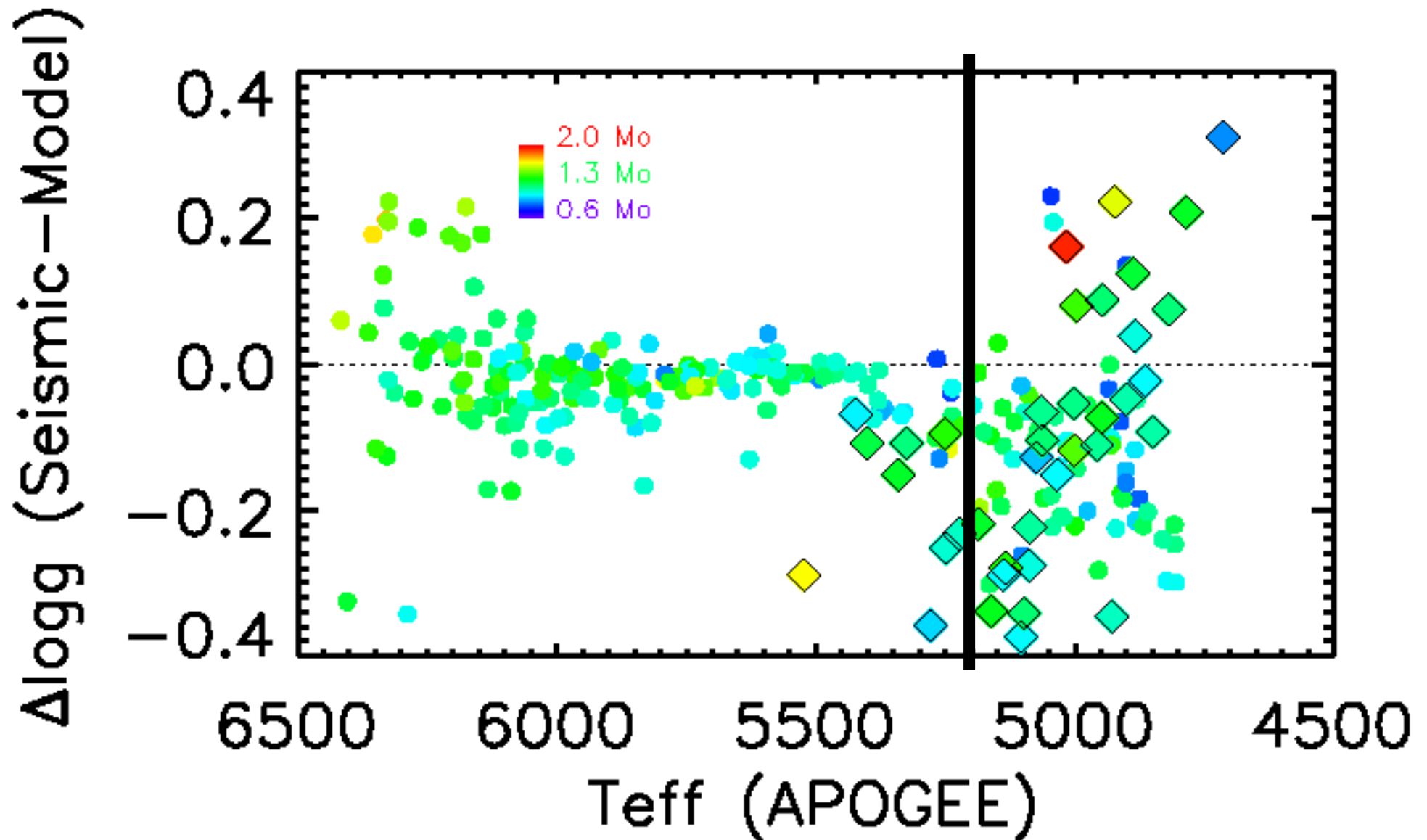
TESS 2-year sky coverage map



APOGEE-Kepler Subgiants



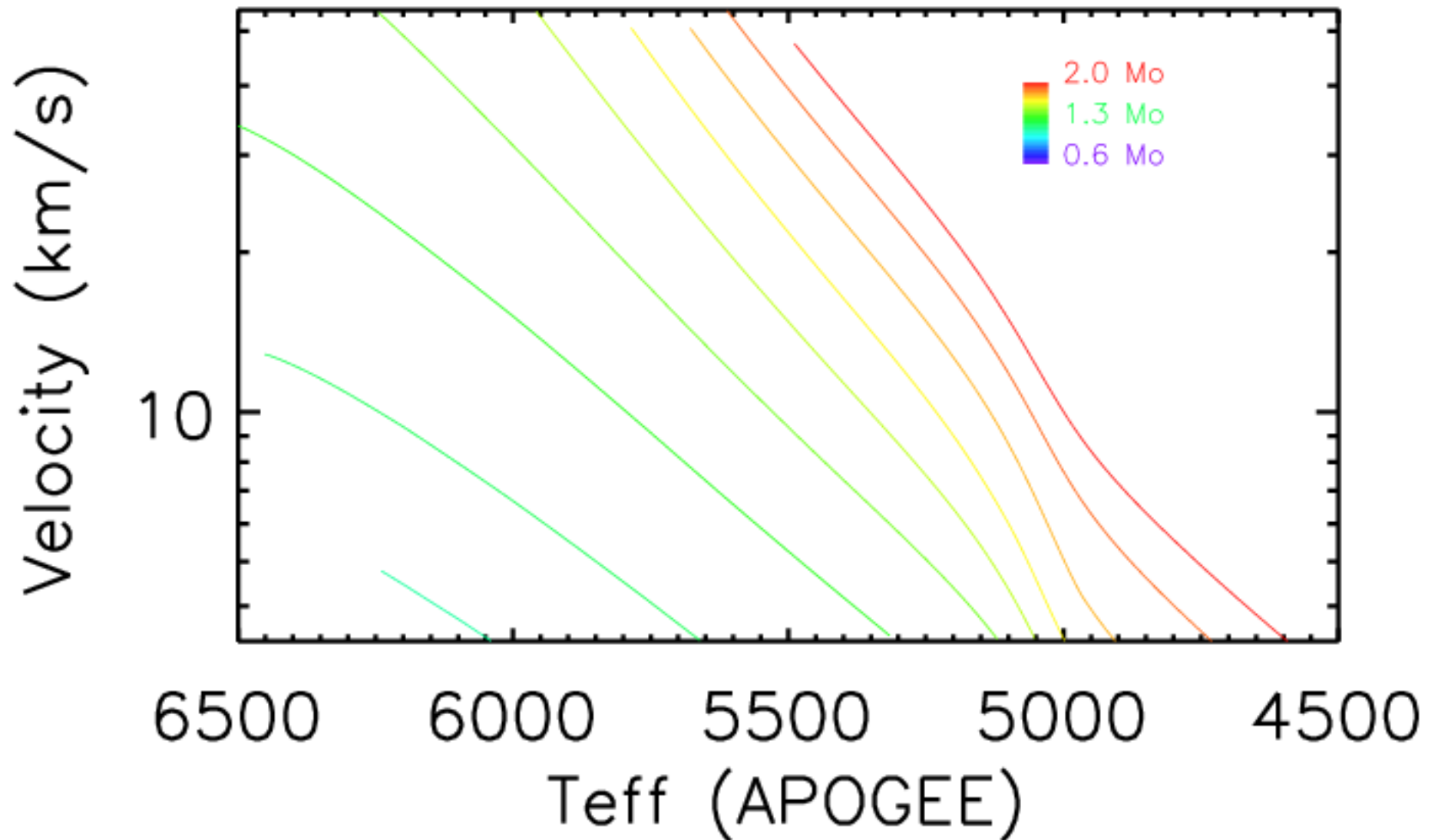
TESS Subgiants



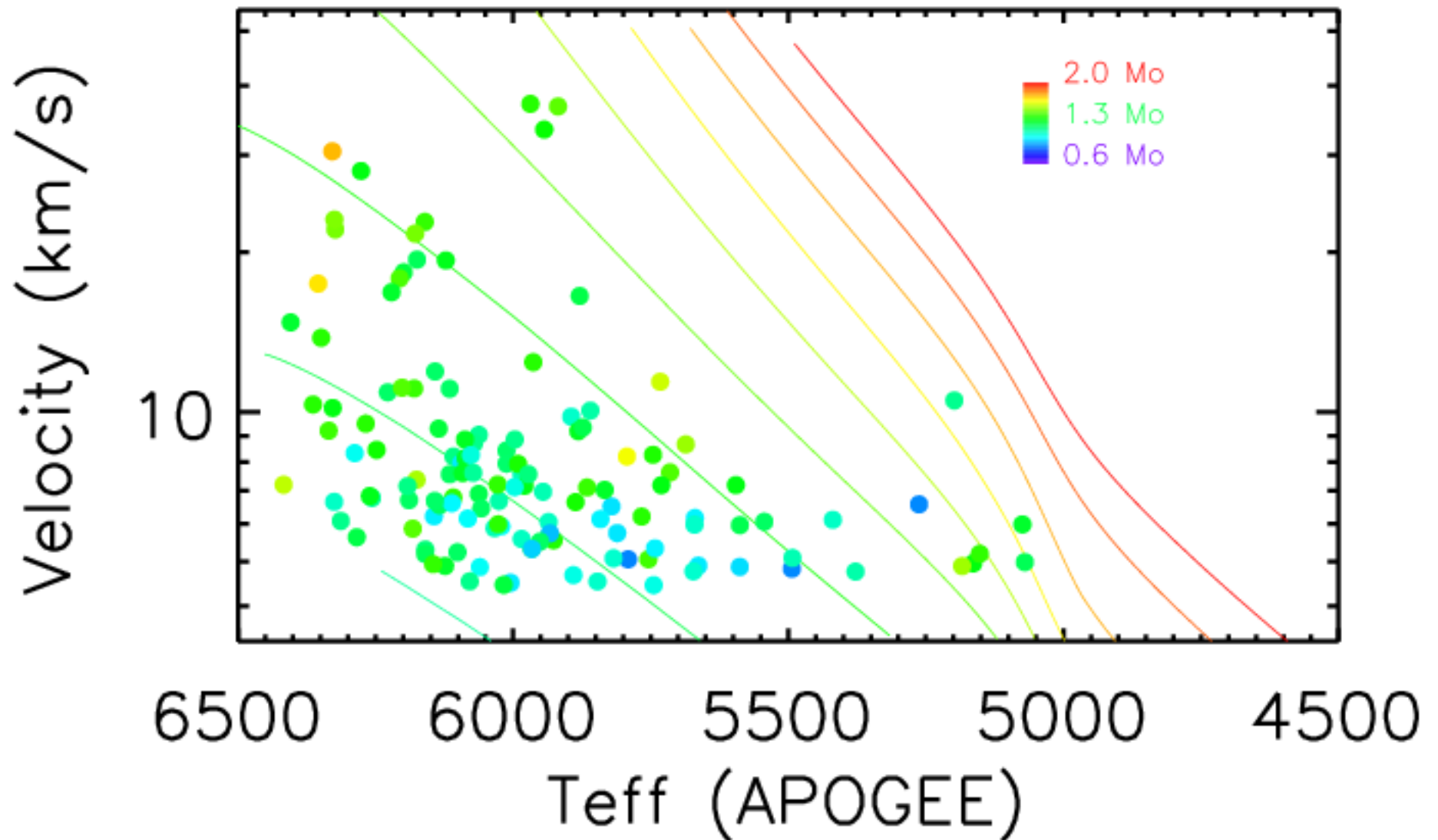
Summary

- Stellar models need calibration
- Model choice impacts inferred mass by $\sim 4\%$
- Predicted HR diagram position \sim correct
- *Kepler* stars are a good but insufficient sample for testing models of subgiants
- *TESS* will allow us to add up to 300 more similarly well characterized subgiants

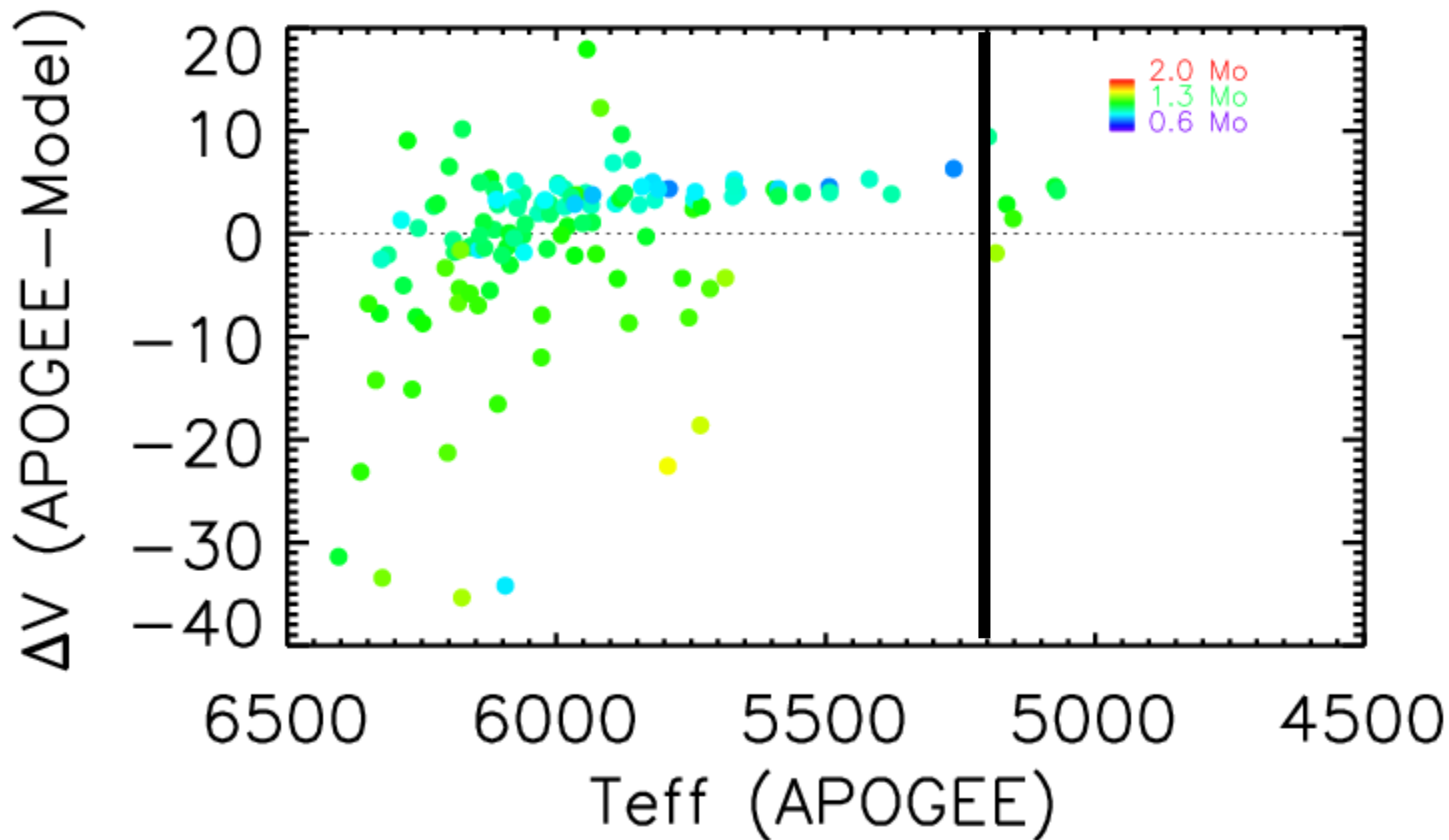
Expected Rotation



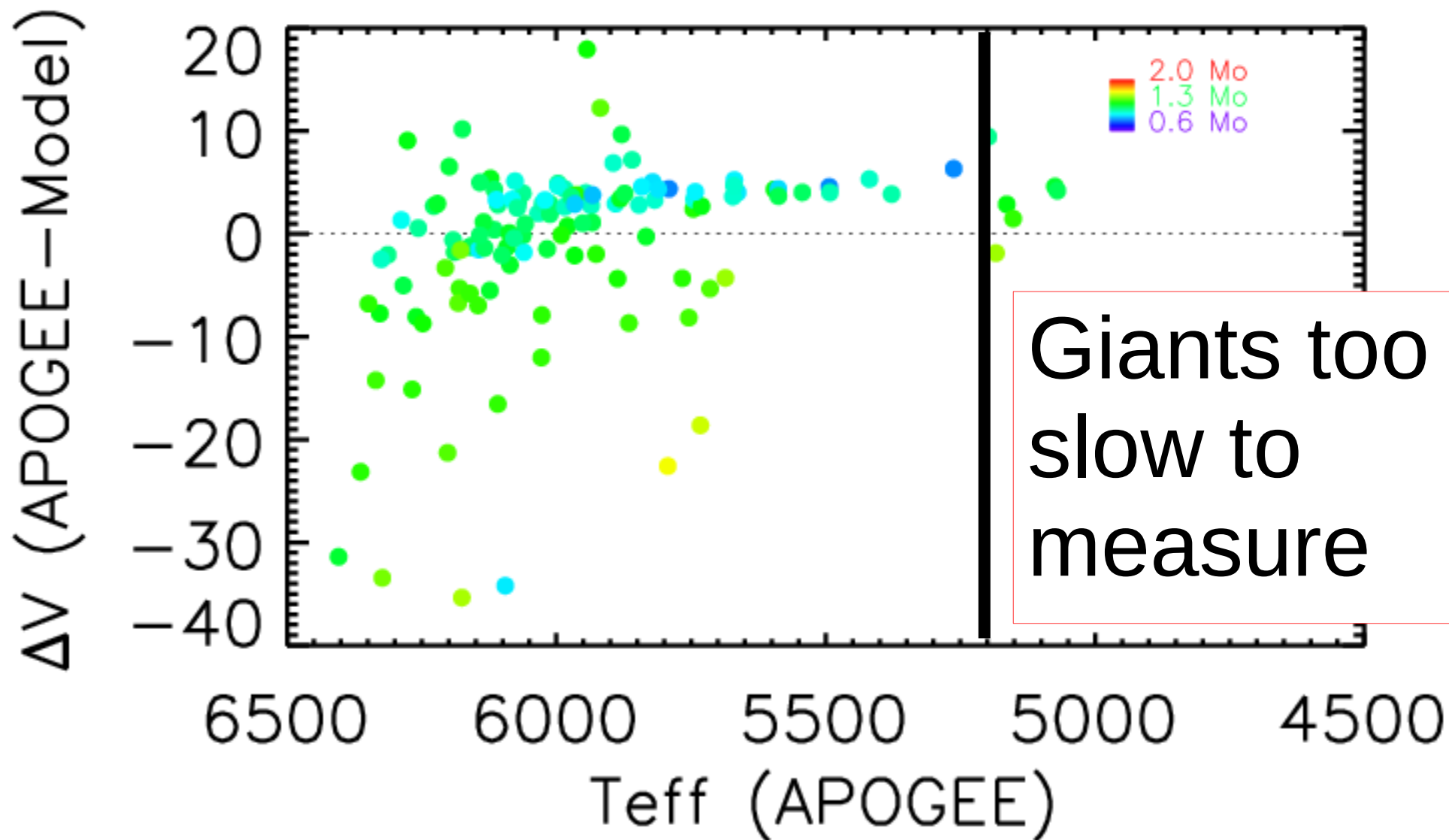
Expected Rotation



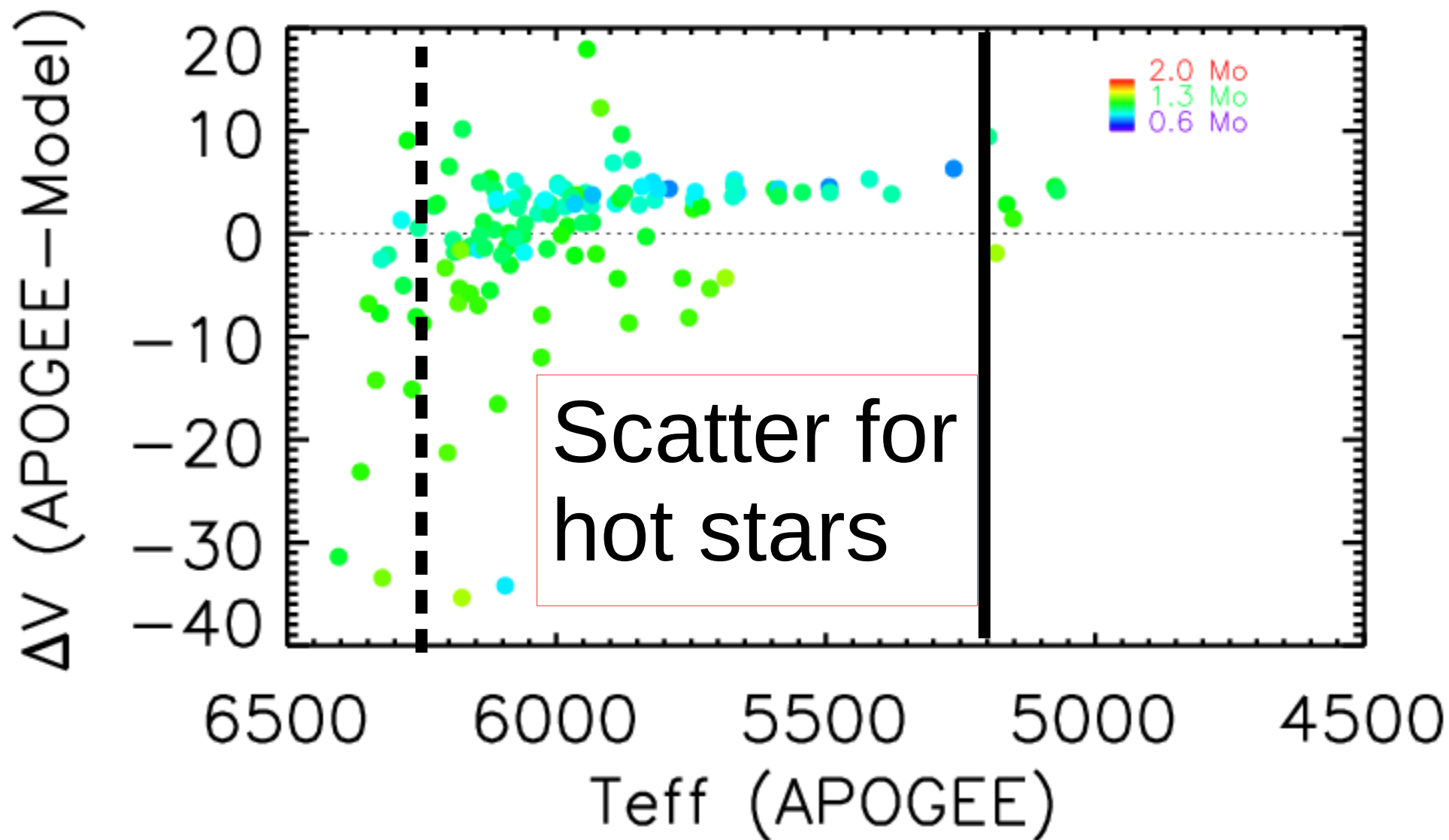
Rotation Difference



Rotation Difference



Rotation Difference



Rotation Difference

